

# Continuous comparison polarimetric receiver for 22, 37/43 and 86 GHz

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# Science motivation

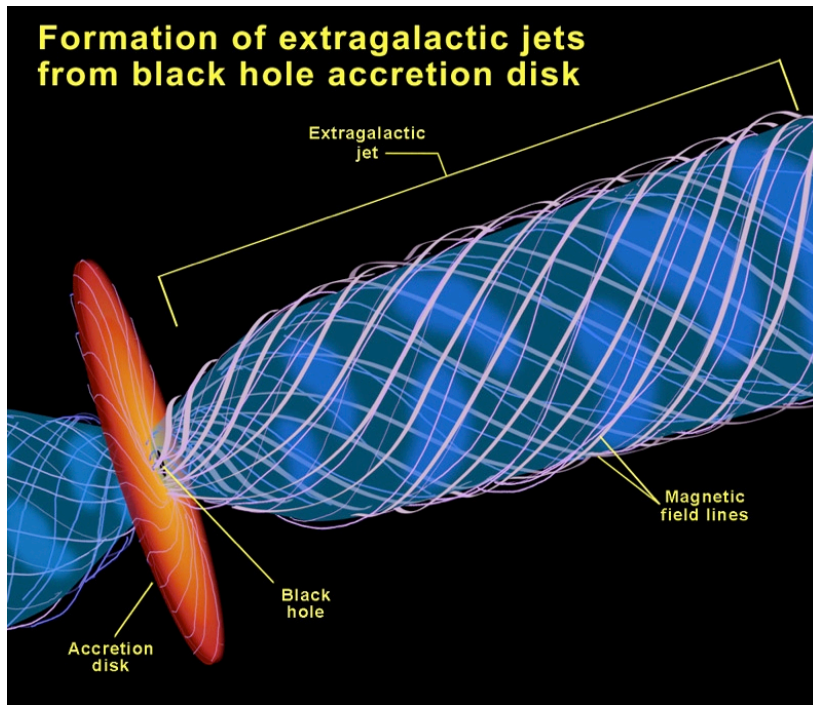
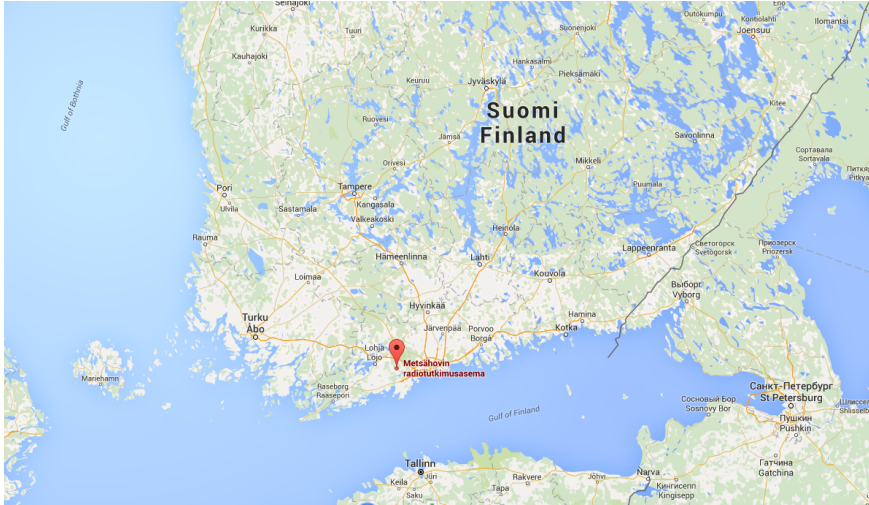


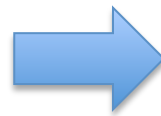
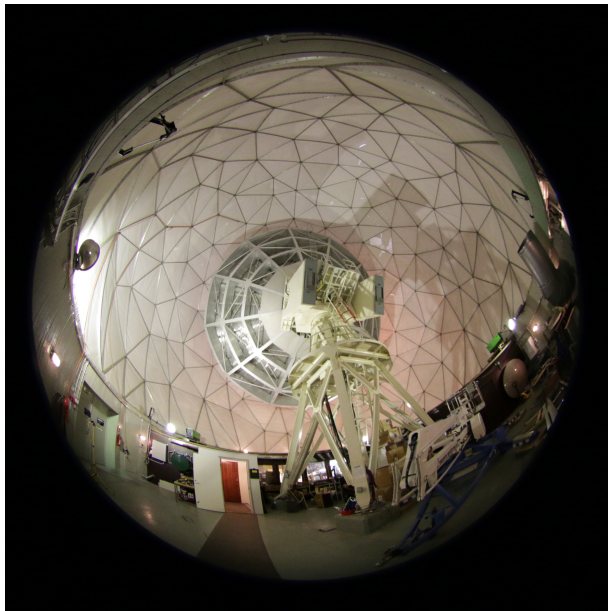
Image credit: NASA / Ann Field

- Jets form through interaction between magnetic fields and black hole / accretion disk
- Polarization is a probe for magnetic fields
- Fast-cadence single-dish observations can help to distinguish ordered and turbulent B-field models
- Circular polarization probes particle composition

# Metsähovi Radio Observatory



- Radome-enclosed 14-m telescope
- Currently most of the observing time is devoted to 22/37 GHz observations of AGN
- Participates in EVN, GMVA and Geodetic VLBI sessions



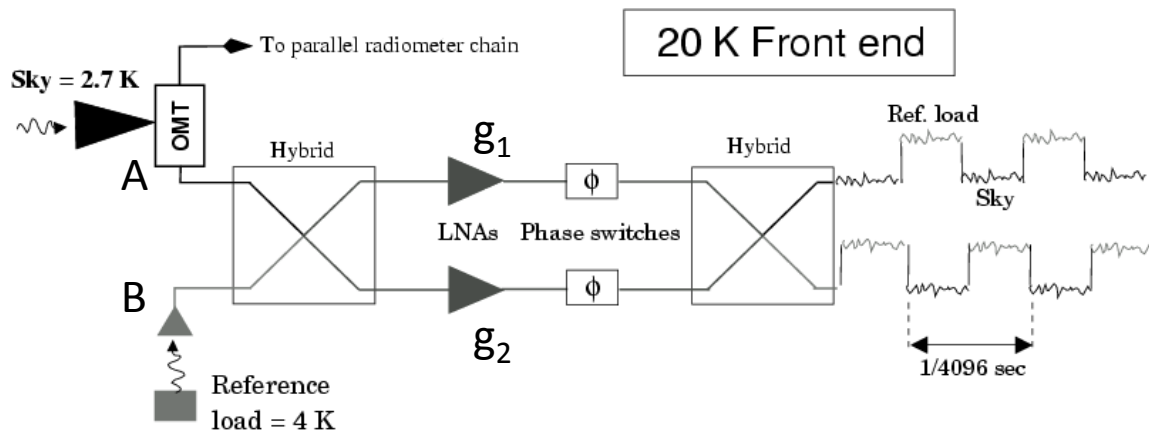
Most beneficial would be to have one receiver that can do everything!

# Technical requirements

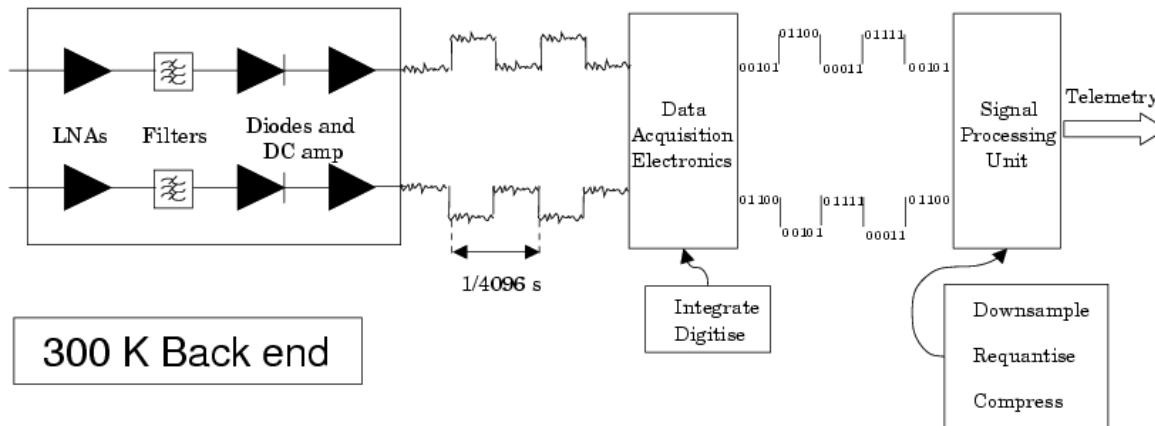
- Polarimetric receiver
  - Both linear and circular
- Multiple frequencies (22 / 37-43 / 86 GHz)
  - Fast switching for VLBI to enable phase transfer
- Capable for single-dish observations as well
  - High sensitivity ( $\sim 10$  mJy source brightness)
  - Reduction of  $1/f$  noise
- Confined to a relatively small space (max diameter 500mm)



# Continuous comparison receiver



Baseline Planck LFI  
Pseudo-correlation receiver  
(Mennella et al. 2003, A&A, 410, 1089)



$$V_l - V_r = \mp s(A^2 - B^2)g_1(t)g_2(t)$$

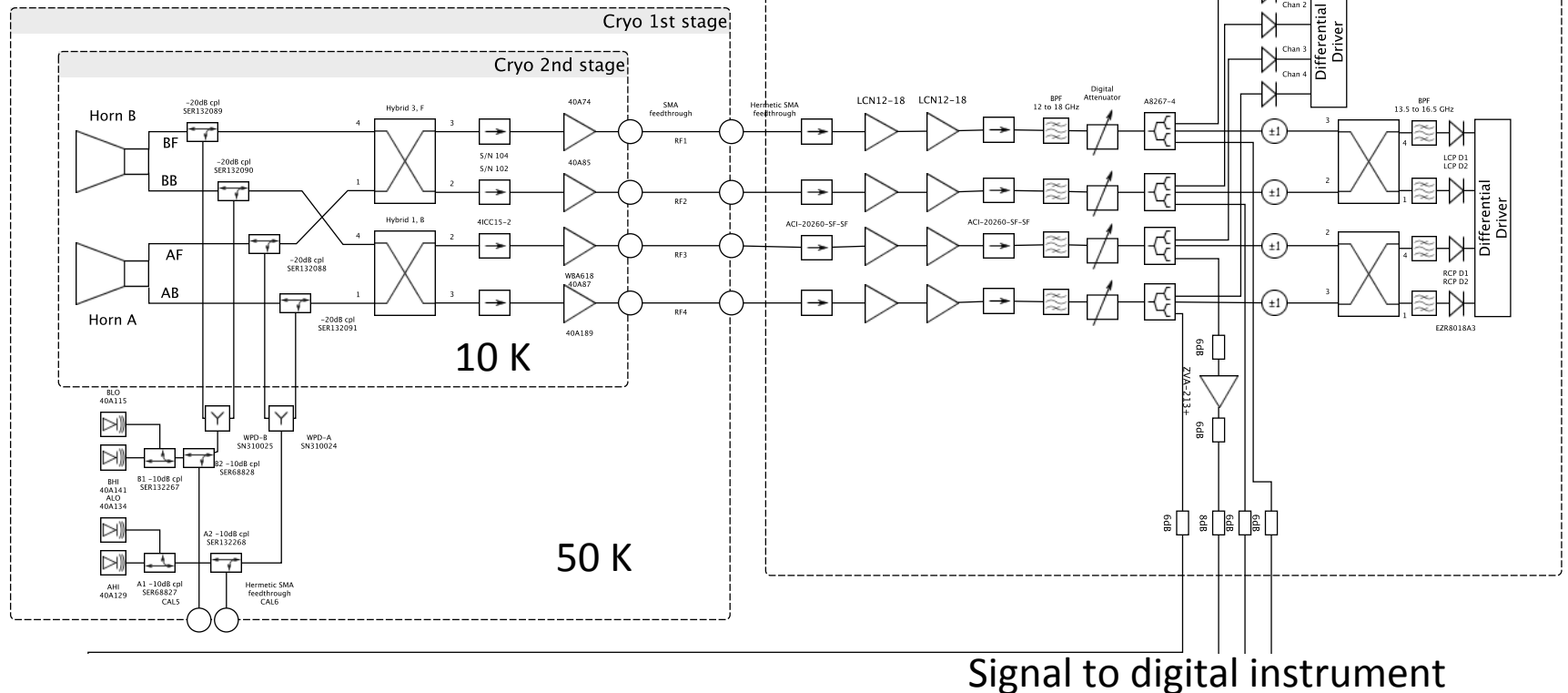
$(A^2 - B^2)$  is much smaller than the sum of the first terms so that the effect on gain variations is smaller

# 12-18 GHz polarization receiver KuPol

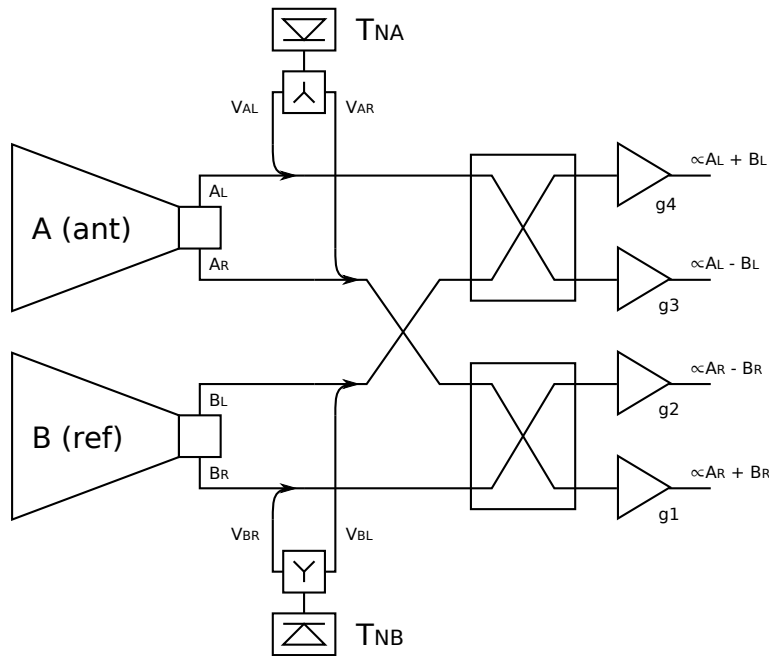
- Mounted on the 40-m telescope at Owens Valley Radio Observatory (Caltech)
- Operational since May 2014
- Hybrid design with both analog and digital backends to reduce commissioning time



Designed by Dr. Oliver King at Caltech

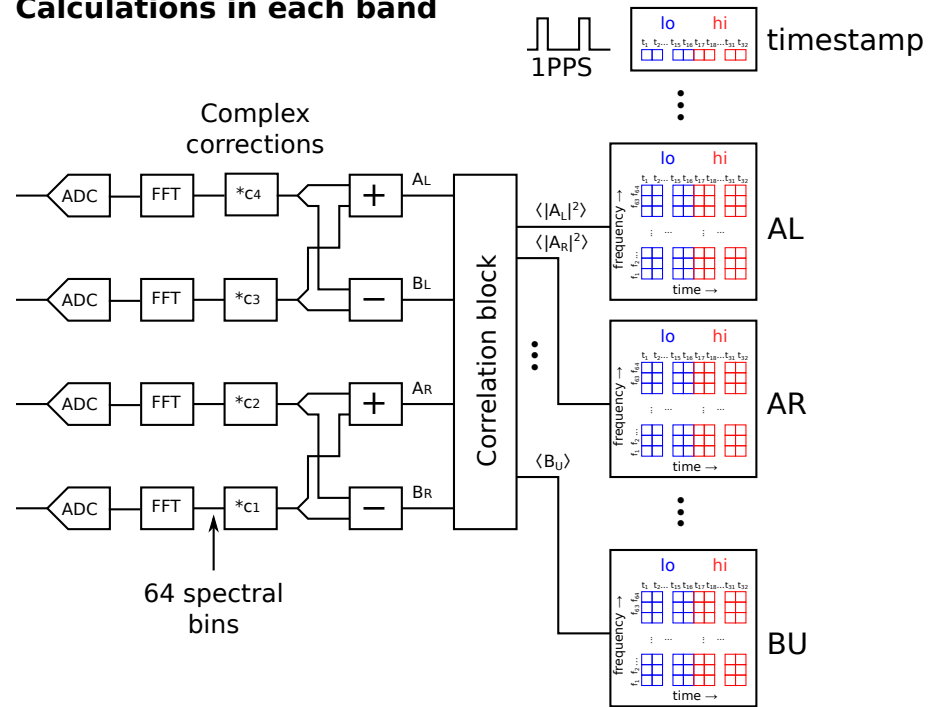


# Polarization through correlation



Split into 12 bands between 12 and 18 GHz

## Calculations in each band



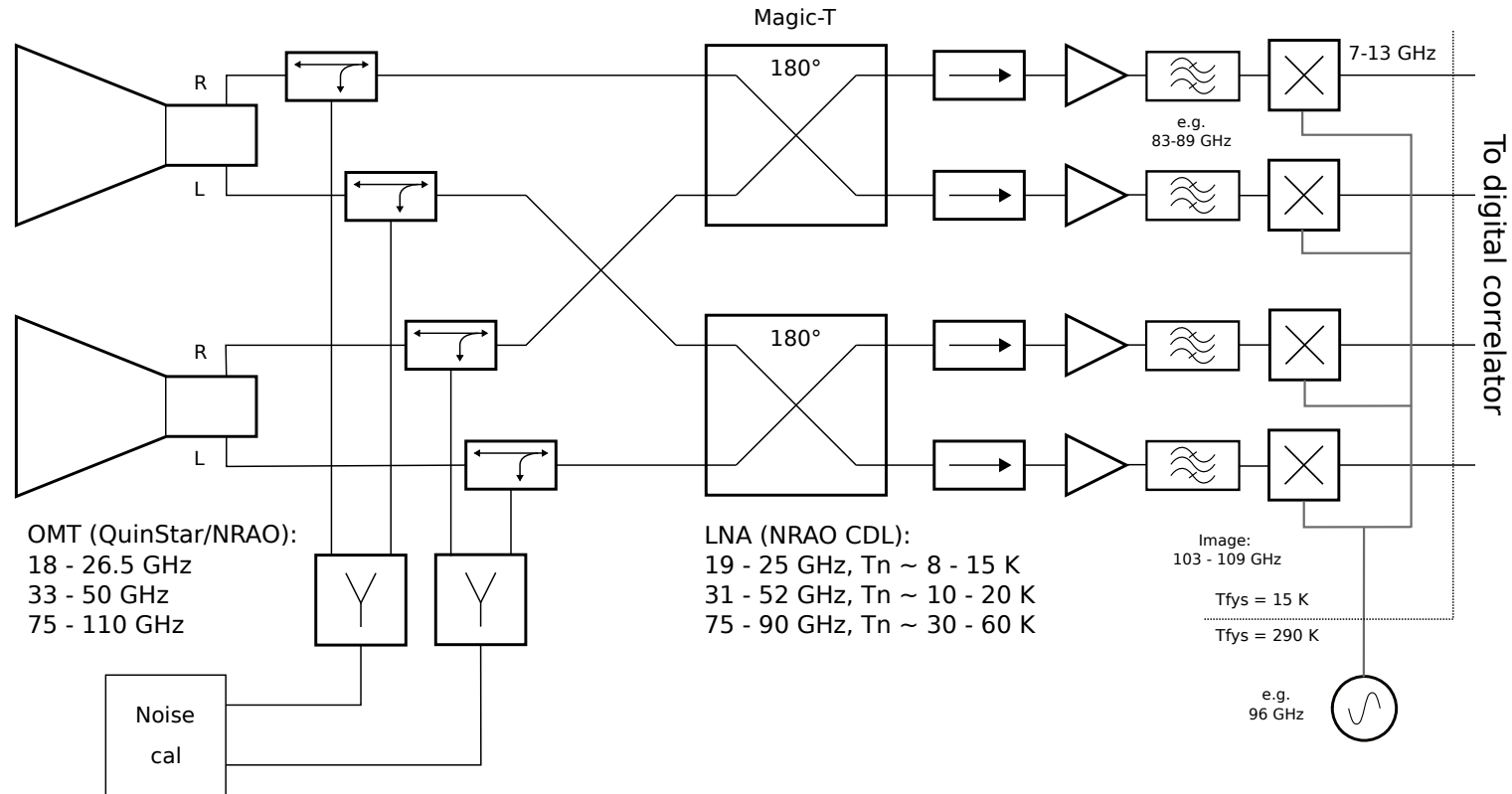
Courtesy of Dr. O. King

Calculations performed with ROACH boards

# Metsähovi frontend

## Metsähovi continuous comparison polarimeter

### Frontend



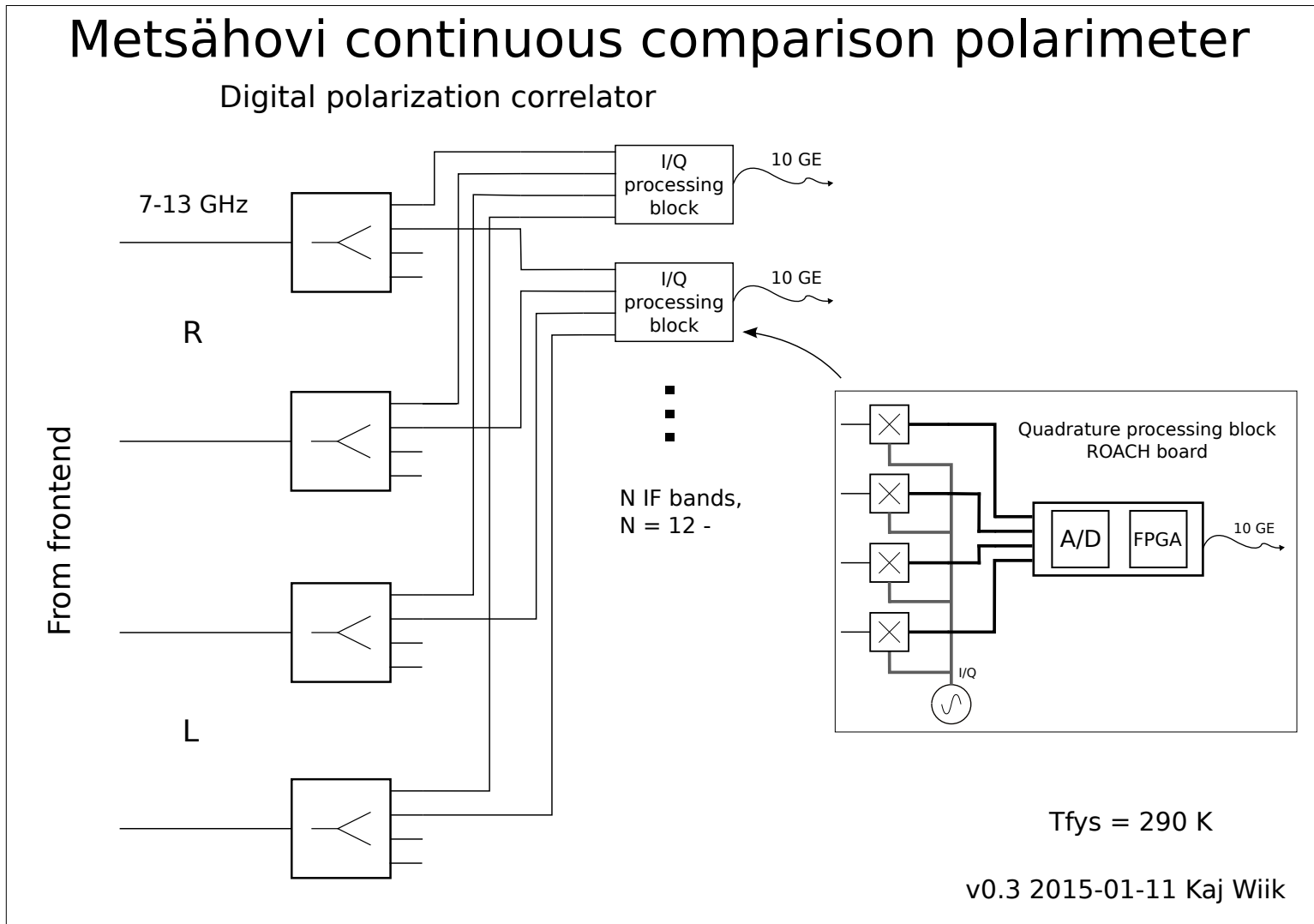
$T_{sys} = 50-80$  K @ 43 GHz

v0.3 2015-01-11 Kaj Wiik

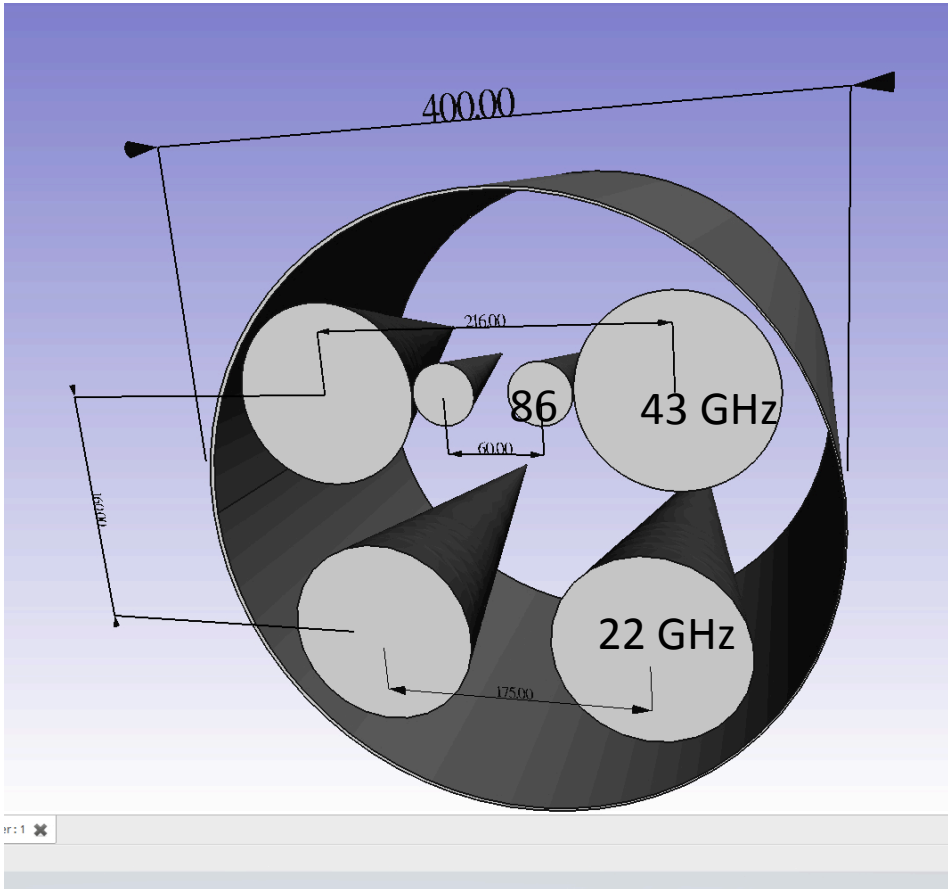
# Digital backend

## Metsähovi continuous comparison polarimeter

### Digital polarization correlator



# Multifrequency setup



- All horns in a single cylinder
  - Switching between frequencies done by changing pointing offsets
    - Between 43-86 GHz takes about 0.9-1.2s depending on elevation
    - Additional stabilization period of (currently) 2s
- ➔ Short enough for phase transfer from 43 to 86 GHz (Middleberg et al. 2003, A&A, 433, 897)

# Rough cost estimate

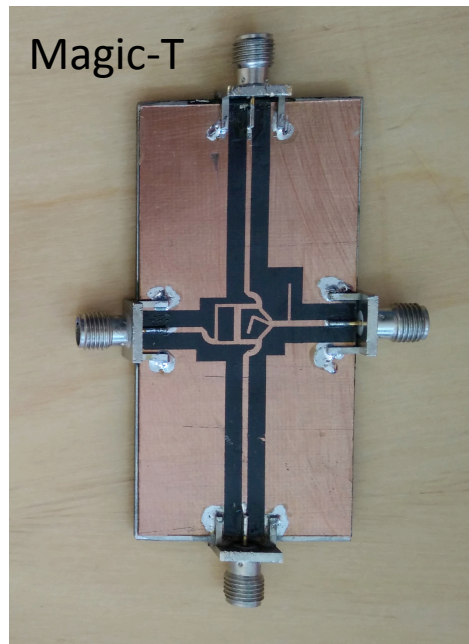
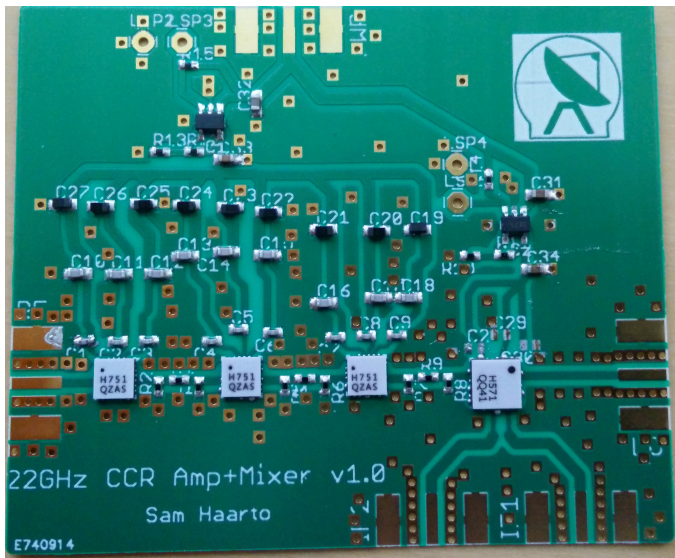
- Cryostat 55 k euro
- Frontends for K / Q / W bands 100k euro each
- Microwave / IF processing 50k euro
- Digital backend 100 k euro

 500 k euro in total (VAT or manpower not included)



# Warm test receiver at 22 GHz

- Master's student project
- Linear feeds
- DBBC for data readout
- Plan to implement a software simulator for digital backend testing



# Summary and outlook

- Need for a receiver that can do both single-dish and VLBI observations
  - Requires high sensitivity ( $\sim 10$  mJy)
- Multiple frequencies and fast switching
- Linear and circular polarization
- ➔ Continuous comparison receiver
  - Start with 37/43 GHz to test concept and add 22 and 86 GHz later
  - Work on a 22 GHz test setup on-going as a Master's student project